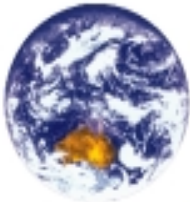


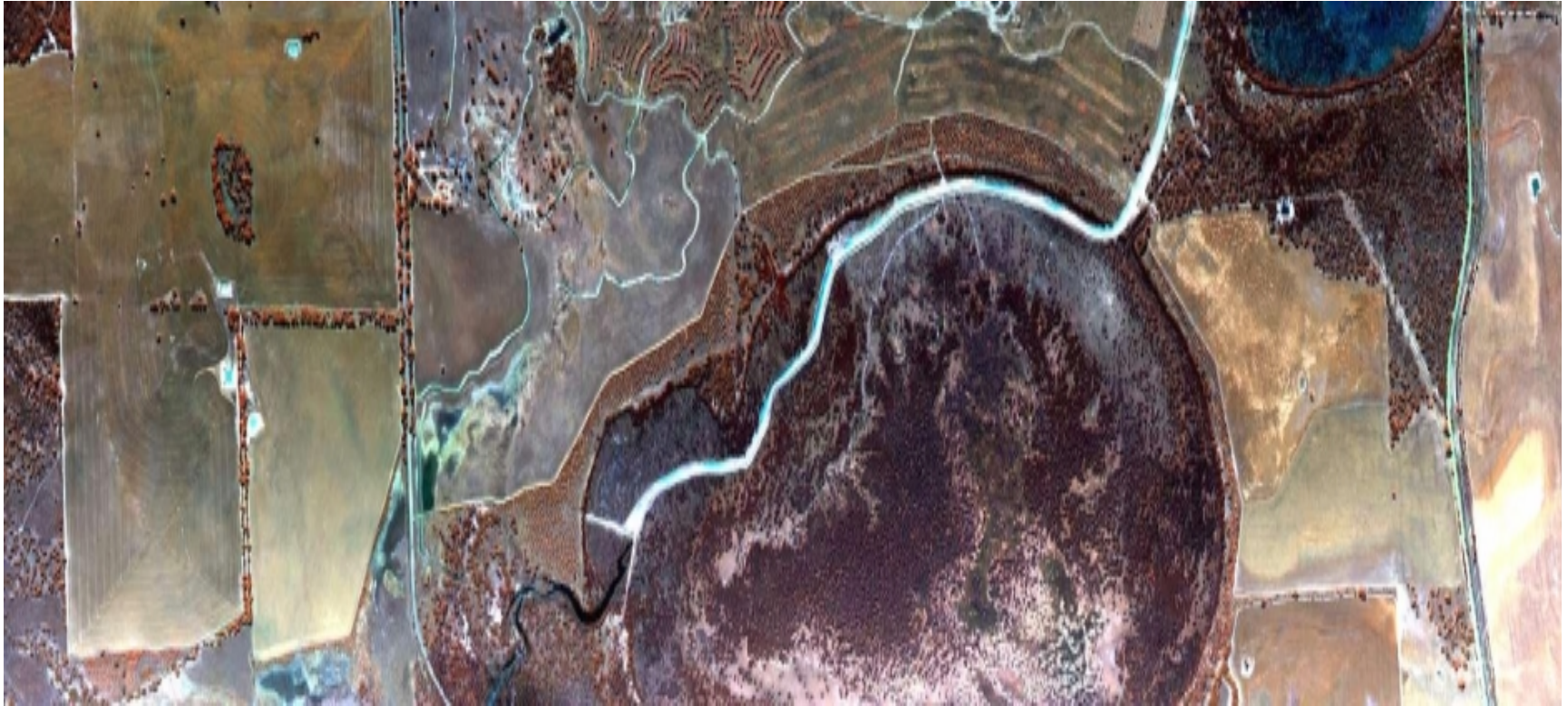


EOC Hyperspectral Science Meeting May 4-5, 1999



**OFFICE OF SPACE SCIENCE & APPLICATIONS
EARTH OBSERVATION CENTRE**

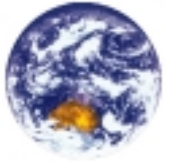
CSIRO Office of Space Science & Applications
Earth Observation Centre



CEM, CLW, CWE, CFFP, CMIS, CMR

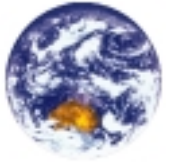
CSIRO Office of Space Science & Applications
Earth Observation Centre

Outline of Discussion



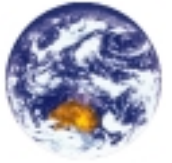
- ❖ Atmospheric Profiles (Sonde Data)
- ❖ Weather Station Data (Q,RH,Ta,V)
- ❖ Spectral Irradiance (Diffuse & Direct)
- ❖ Target Reflectance (PIFs and Panels)
- ❖ Atmospheric Modelling & Correction

Sonde or Atmospheric Profile Data

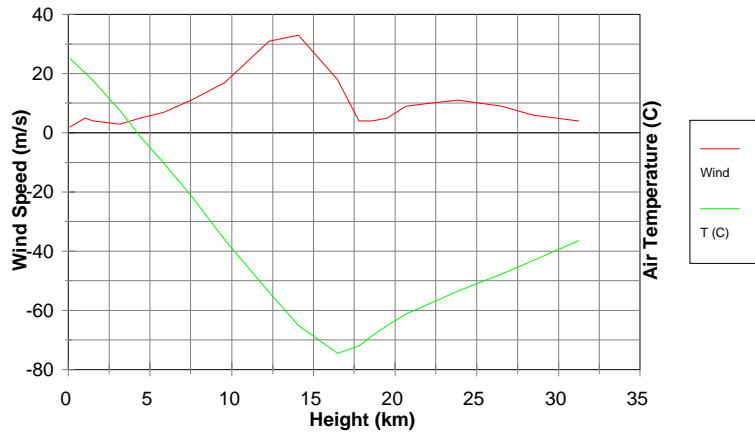


- ❖ Profile data are: $[P, T_a, RH, [V]](z)$
- ❖ Or Pressure, Air Temperature, Humidity, Wind Speed as a function of altitude (z)
- ❖ This is used in atmospheric correction and possibly surface energy balance
- ❖ Wind speed is not always available

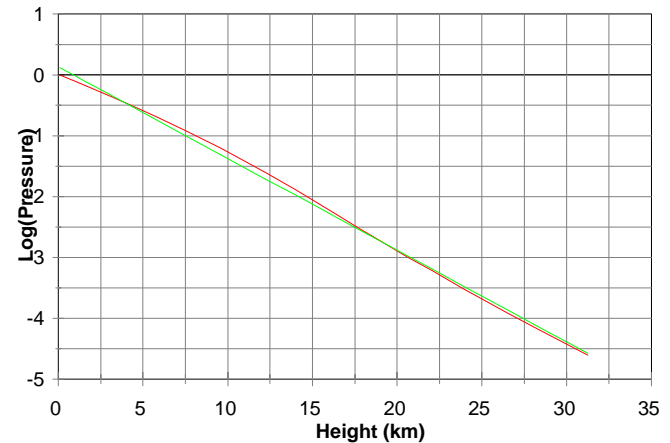
Alice Springs Upper Air Statistics (UAS)



Alice Springs UAS November
Upper Air Winds



Alice Springs November
LogP vs Altitude



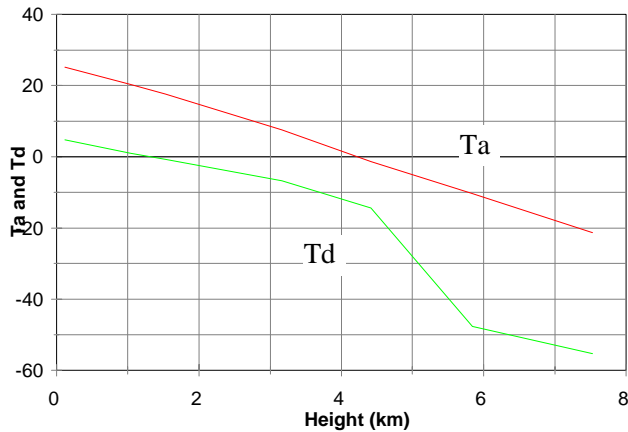


- ❖ Upper Air Statistics (Maher & Lee, 1976)
- ❖ Nearby airport/met station
- ❖ EOC/CAR Sondes
- ❖ Aircraft measurements (esp at various heights)
- ❖ (Water Vapour can be RH%, MR, vapour pressure, dew point (Td) or vapour density)

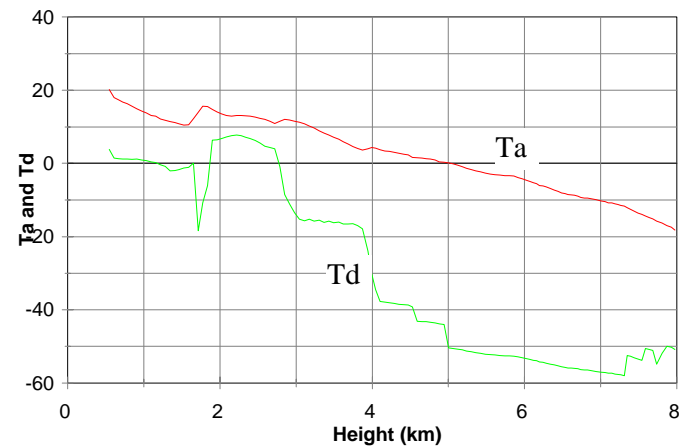
Alice Springs Average UAS vs Local Radiosonde



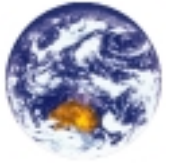
Alice Springs UAS November
Air Temp vs Dew Point



Alice Springs November 28, 1998
Air Temp vs Dew Point



Simplified Profile Models



$$\rho_a(z) = \rho_a(0) e^{-c_1 z}$$

$$\rho_w(z) = \rho_w(0) e^{-c_2 z}$$

$$\frac{P(z) \left(\frac{T(0)}{T(z)} \right)^{0.5}}{P(0)} = e^{-c_3 z}$$

$$P(z) = P(0) e^{-c_1 z}$$

where:

P is total pressure (Pa)

T is absolute temperature (K)

ρ_a is the density of air

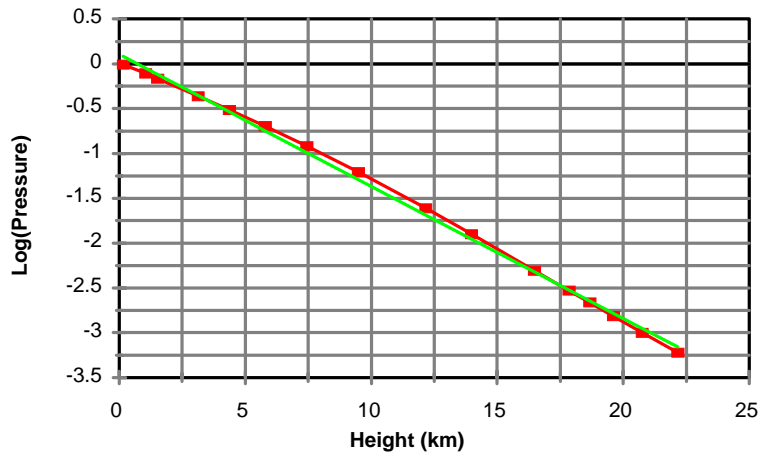
ρ_w is water vapour density and

z is altitude above the surface.

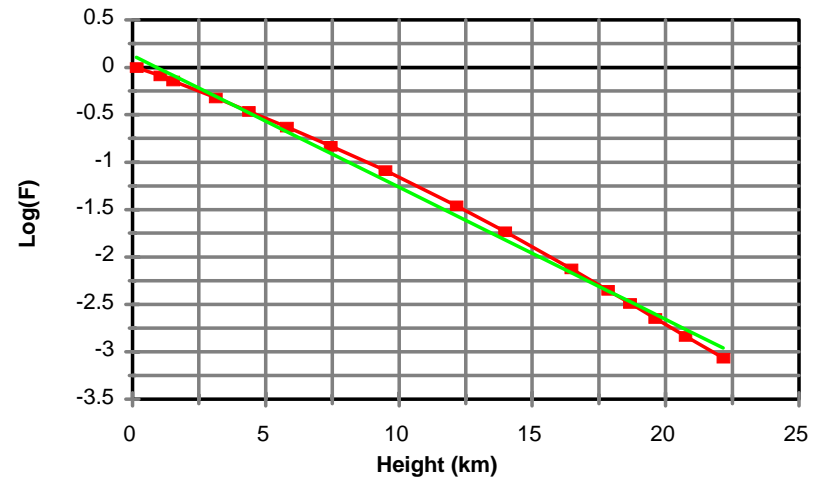
The constants c_1 & c_3 - Air Pressure and Temperature



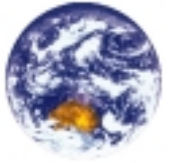
Wagga March 10 yr Atmosphere
LogP vs Altitude



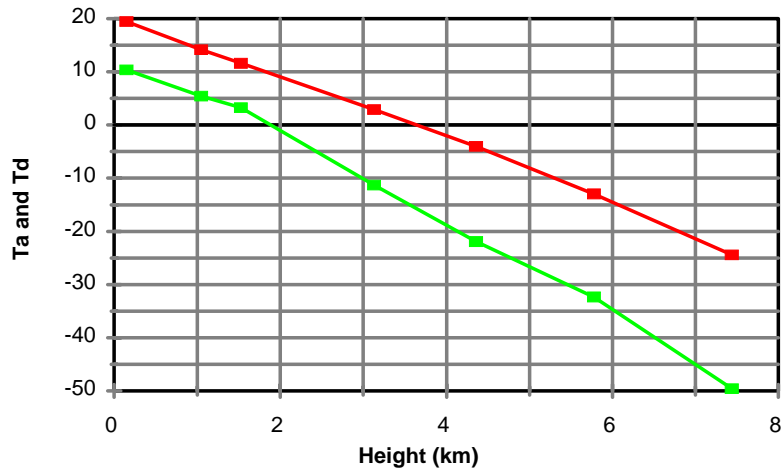
Wagga March 10 yr Atmosphere
Log(F) vs Altitude



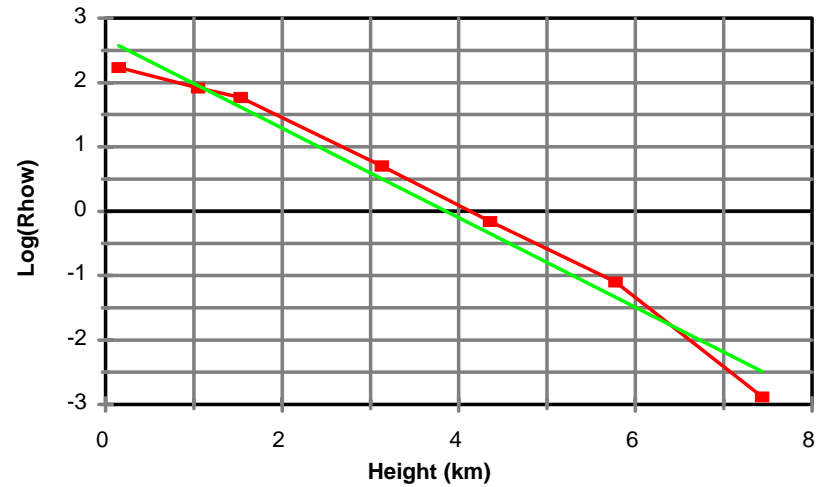
Atmospheric Water Vapour - c2

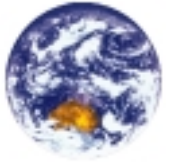


Wagga March 10 yr Atmosphere
Air Temp & Dew Point



Wagga March 10 yr Atmosphere
LogRhow vs Altitude

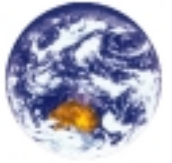




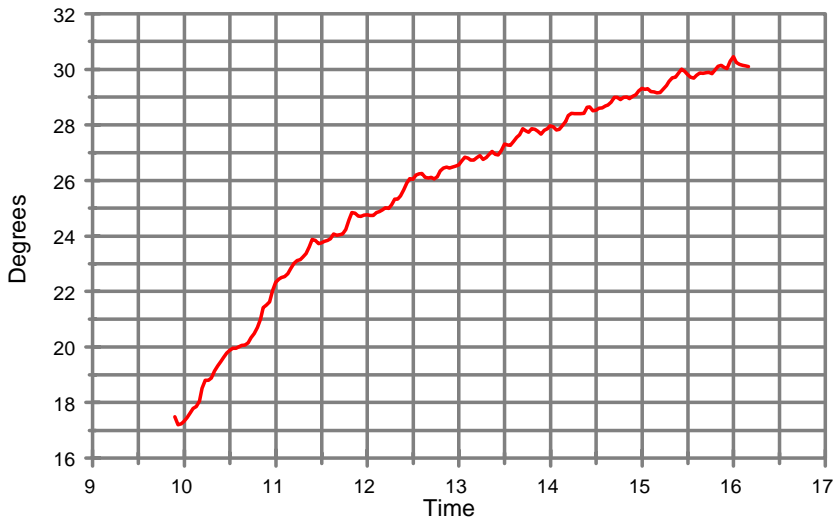
- ❖ Combinations of UAS and local information are usually needed to construct a profile
- ❖ Simplified models allow the surface (Met Station) data to drive the atmospheric profile
- ❖ The detailed shape of the profile is not critical

Weather Station Data

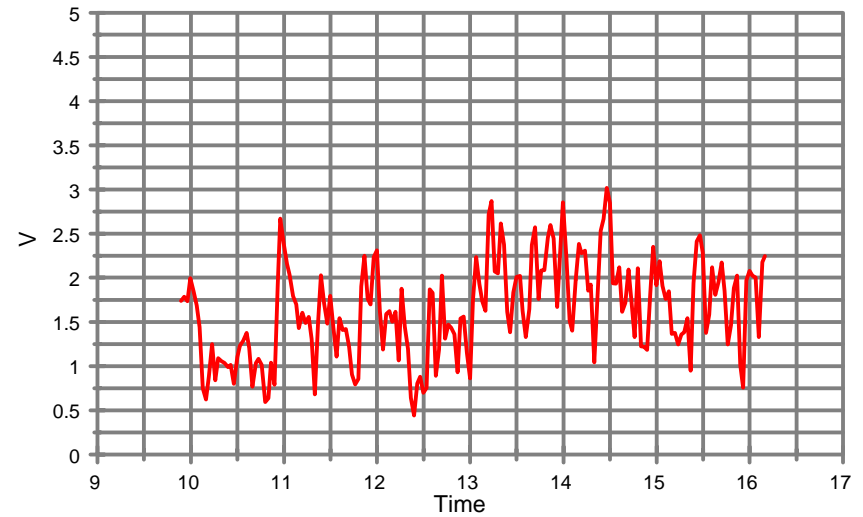
Air Temperature & Wind Speed

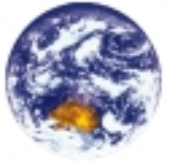


Dickson Oval 10 March 1998
Air Temperature



Dickson Oval 10 March 1998
Wind Speed





Atmospheric Water Vapour

$$\begin{aligned} W(h) &= \int_0^h \rho_w(z) dz \\ &= \frac{\rho_w(0)}{c_2} (1 - e^{-c_2 h}) \\ &= W (1 - e^{-c_2 h}) \end{aligned}$$

$$RH = 100 \times \frac{\rho_w}{\rho_w^*(T)}$$

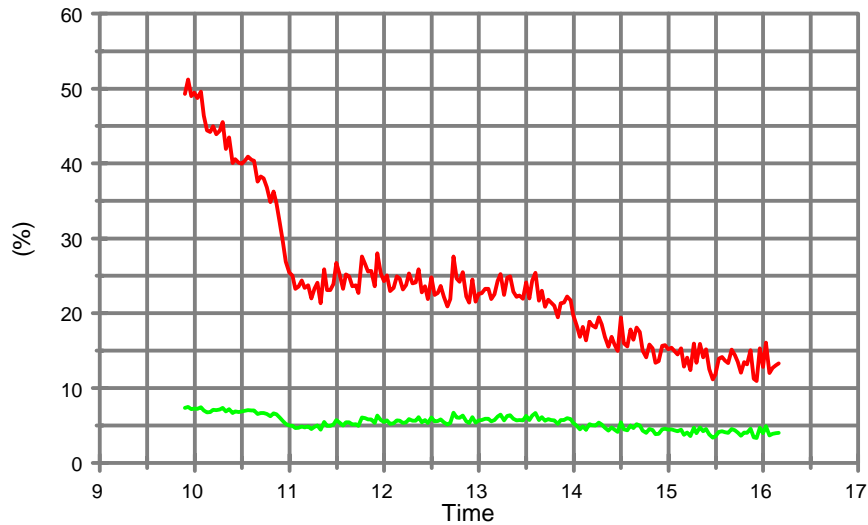
$$\rho_w^* = A e^{a_0 + a_1 A + a_2 A^2} \quad (A = T_0 / T)$$

Weather Station Data

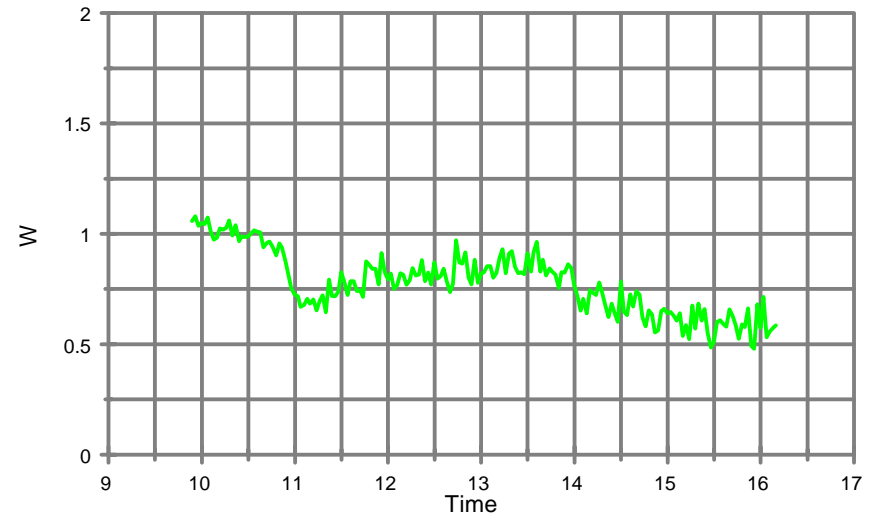
Humidity, ρ_w & W



Dickson Oval 10 March 1998
Humidity & Density of Water Vapour



Dickson Oval 10 March 1998
Precip Water (W)



Solar Radiation Model



$$Q_s = Q'_0 \cos \theta_s t^{m^\alpha / 2}$$

where:

Q'_0 is the exoatmospheric irradiance modified for sun-earth distance

θ_s is the solar zenith angle

m is the air mass (Kaasten's formula)

t is atmospheric transmittance, and

α is a parameter modifying the air mass

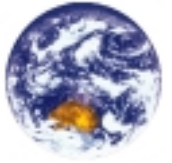
Clear Day Solar Radiation



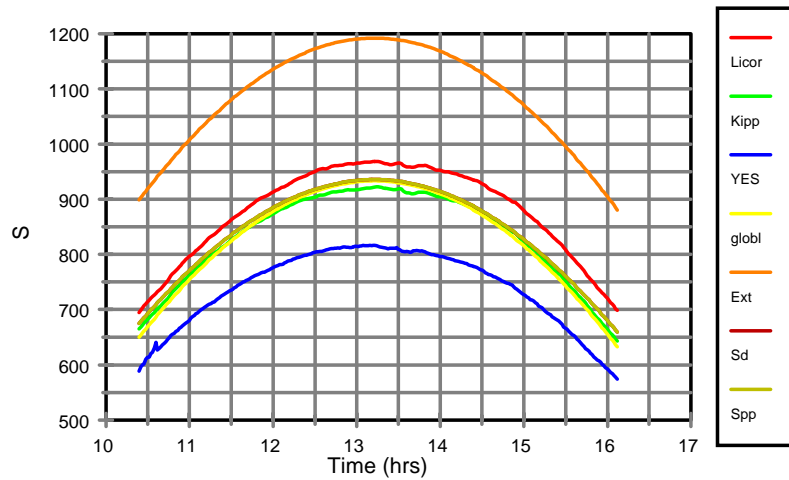
Case	t	α
Prata & Grant ¹ (both)	0.6188	0.450
Prata & Grant (Darnell) ²	0.6188	0.620
Paltridge & Platt ³ (both)	0.6422	0.5939
Paltridge & Platt (Darnell)	0.6469	0.6644
Dickson Oval ⁴ (both)	0.6288	0.5968
Dickson Oval (Darnell)	0.6318	0.6408

1. 248 apparently clear days at Uardry
2. $\alpha = 1.1 + 1.0 \text{ Log } t$
3. 5 years of clear days at Aspendale
4. One day March 3, 1998 at Dickson Oval

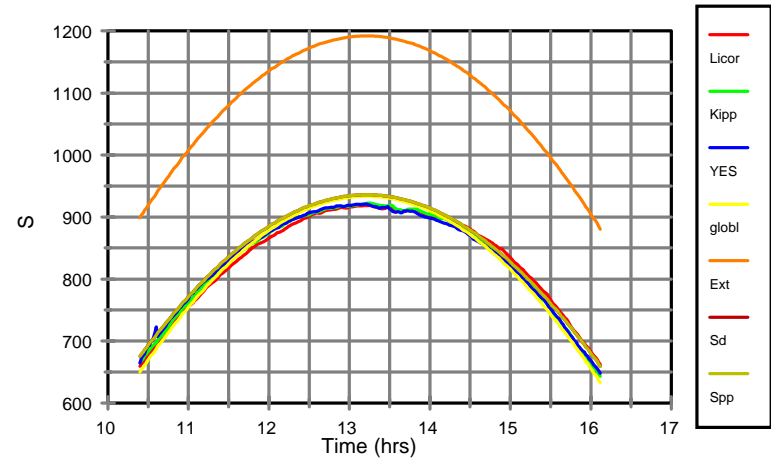
Cross-Calibrating Solar Radiation



Test for Solar Radiation
Licor, Kipp & YES



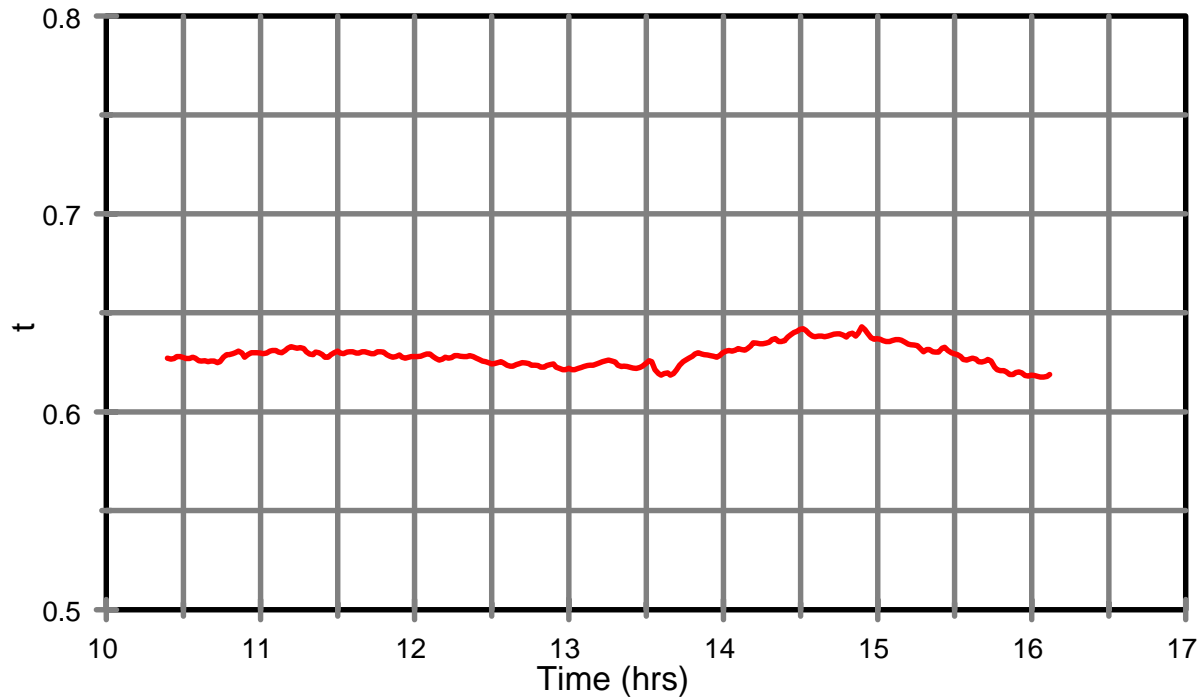
Test for S
Licor, Kipp & YES





Dickson Oval Transmittance

Test for S
Kipp Transmittance

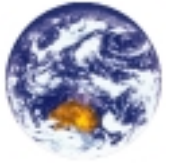


Solar Radiation Data



- ❖ These data provide data integrity and environmental conditions
- ❖ Learning to interpret Kipp transmittance will help “calibrate” atmospheric models
- ❖ Cross calibration with the Yankee MFR

Direct & Diffuse Spectral Irradiance



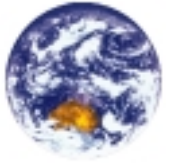
$$E_T(\mu_s, h, \lambda) = \frac{E_T(\mu_s, 0, \lambda)}{T(\mu_s, h, \lambda)}$$

$$E_{dir}(h, \lambda) = \frac{t(\mu_s, \infty, \lambda)}{t(\mu_s, h, \lambda)} \mu_s E_0'(\lambda)$$

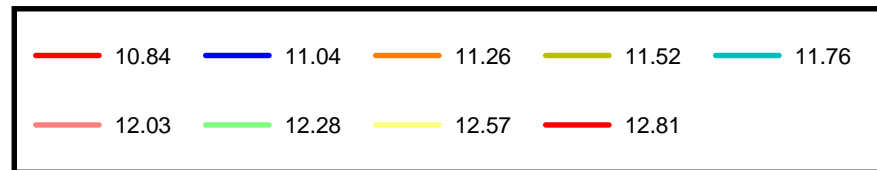
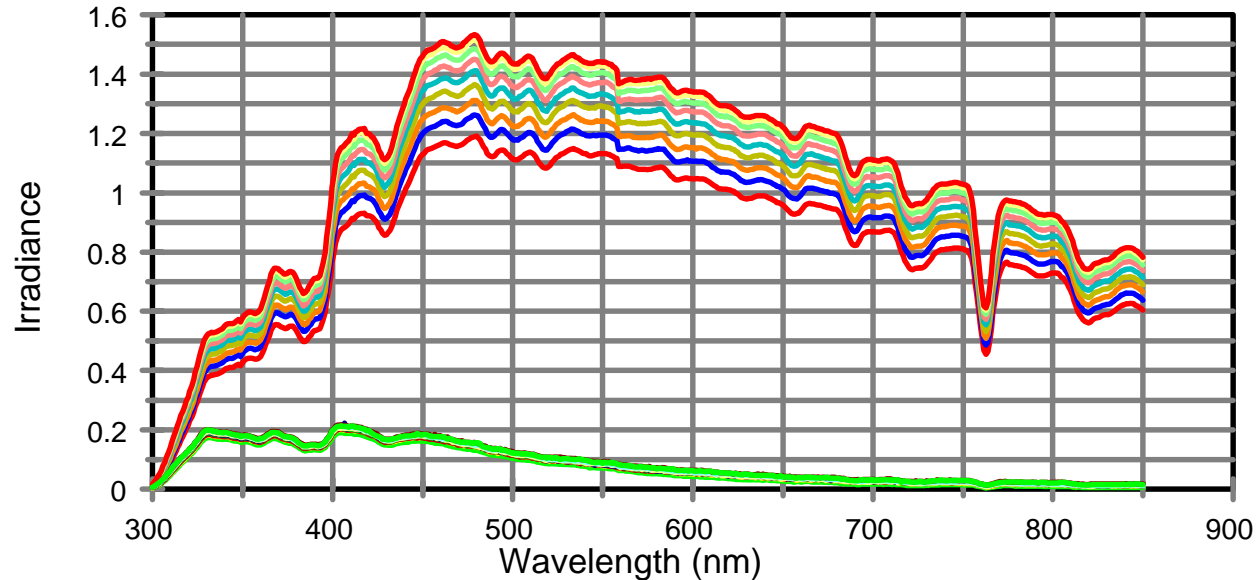
$$E_T^*(\mu_s, h, \lambda) = \frac{E_T(\mu_s, h, \lambda)}{1 - s(h) \rho^*}$$

$$E_{dif}(h, \lambda) = E_T^*(h, \lambda) - E_{dir}(h, \lambda)$$

LiCor Measured Total & Diffuse Irradiance



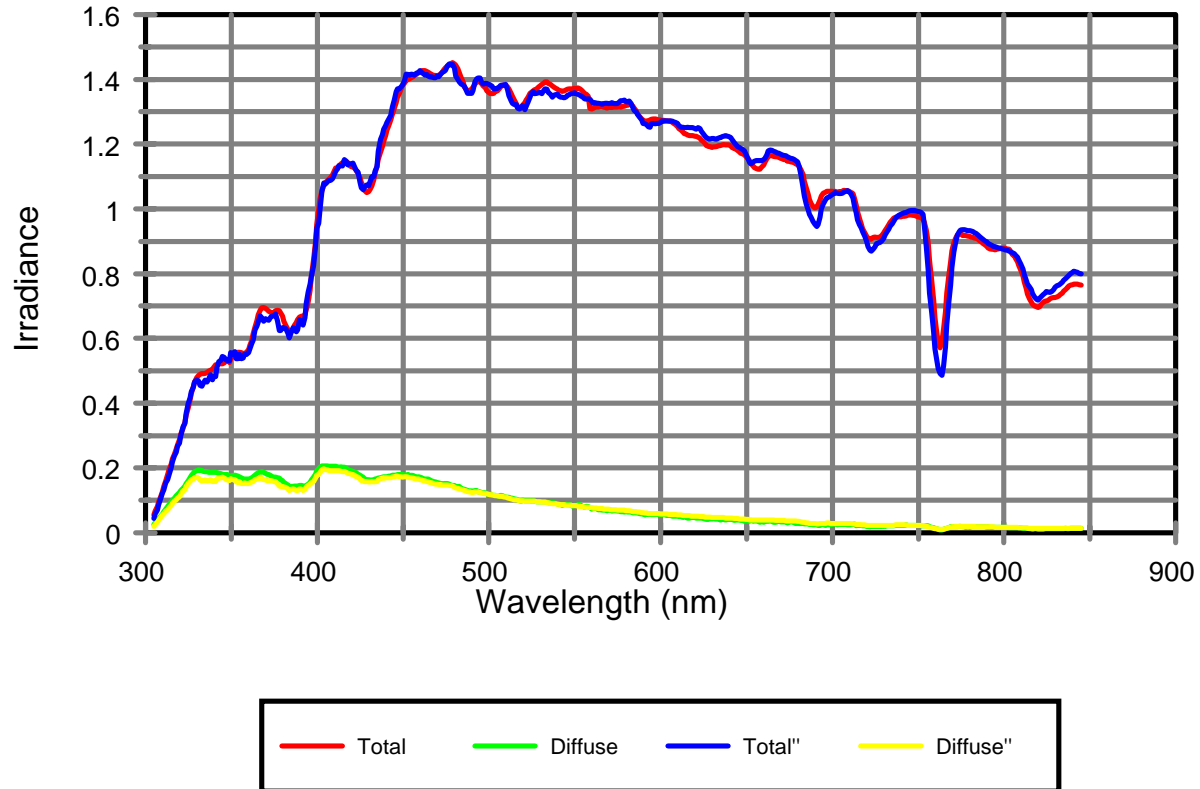
Dickson Oval
LiCor Irradiances (AM)



Inversion of Irradiance Model provides input to Atmospheric Model



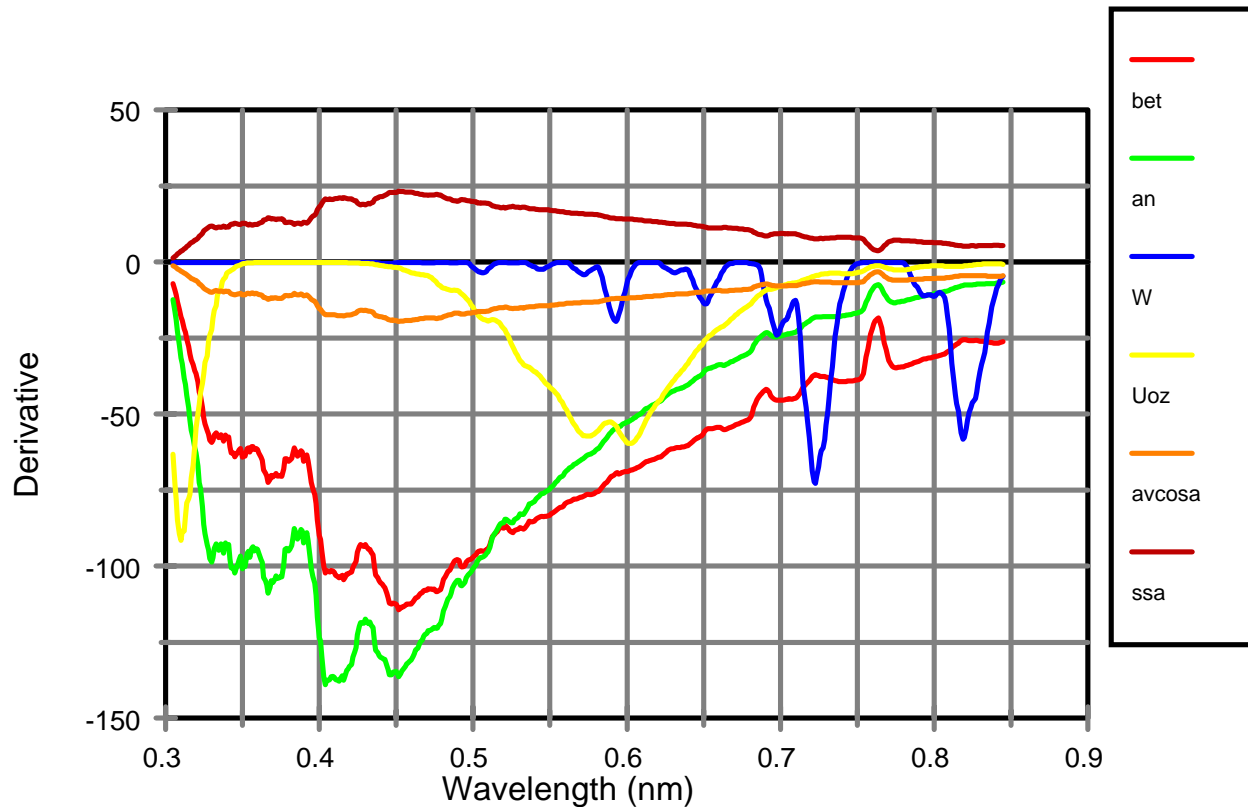
Dickson Oval
LiCor Irradiances (AM)

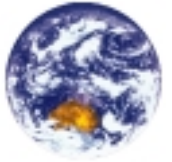


Sensitivity Analysis of Irradiance Model



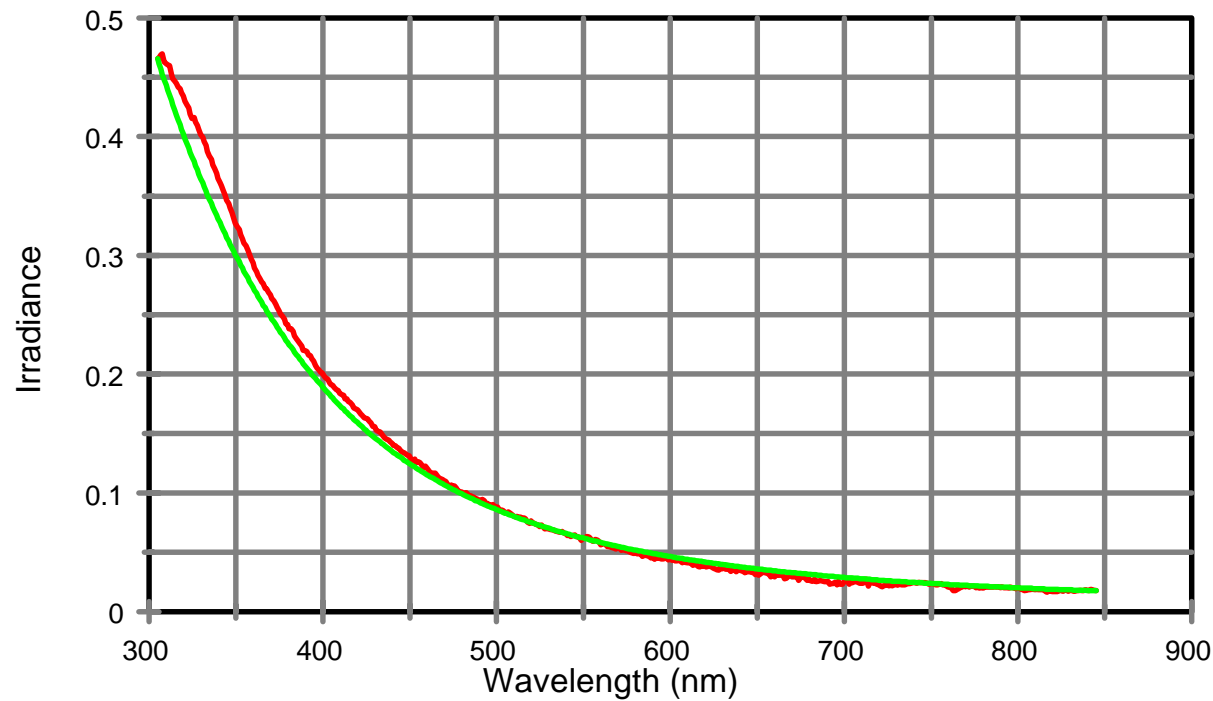
Dickson Oval
Parameter Sensitivities





Diffuse to Total Ratio

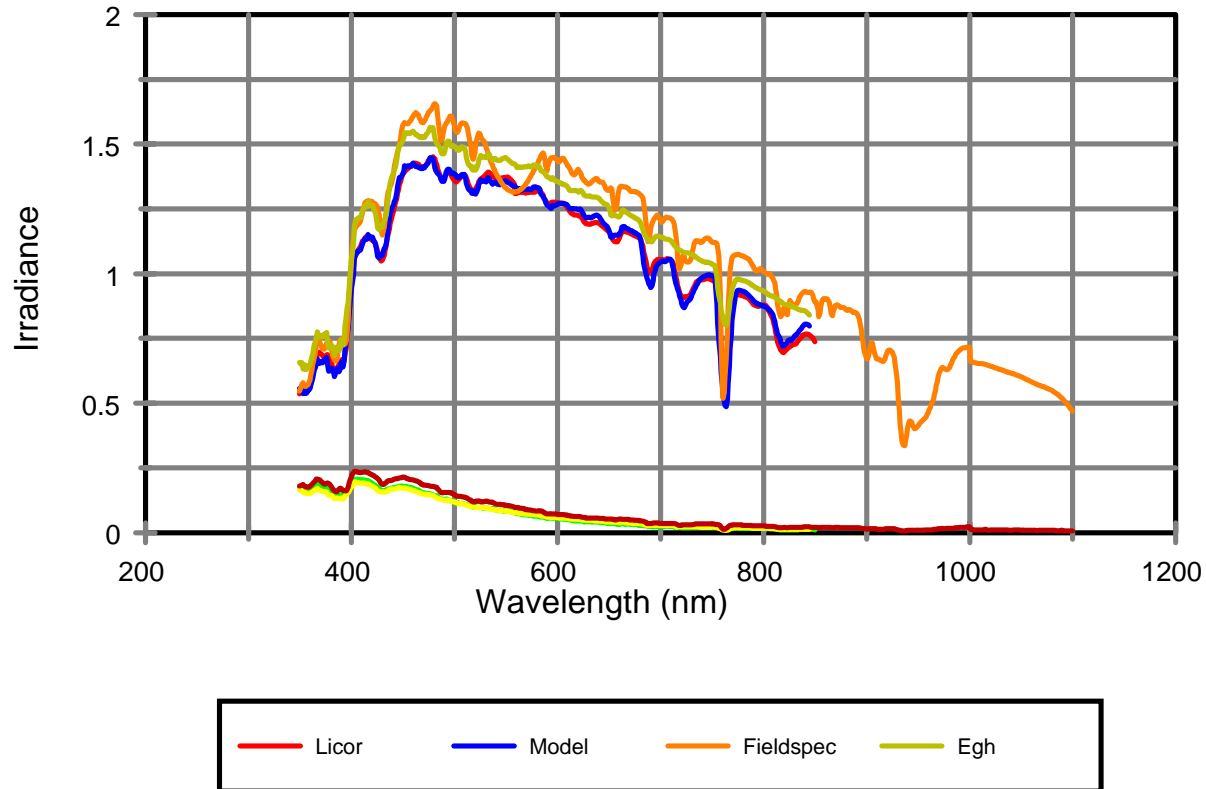
Dickson Oval
LiCor Irradiances (AM)



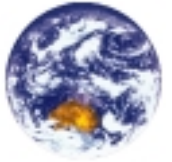
ADS FieldSpec Irradiance



Dickson Oval
LiCor & FieldSpec Irradiances



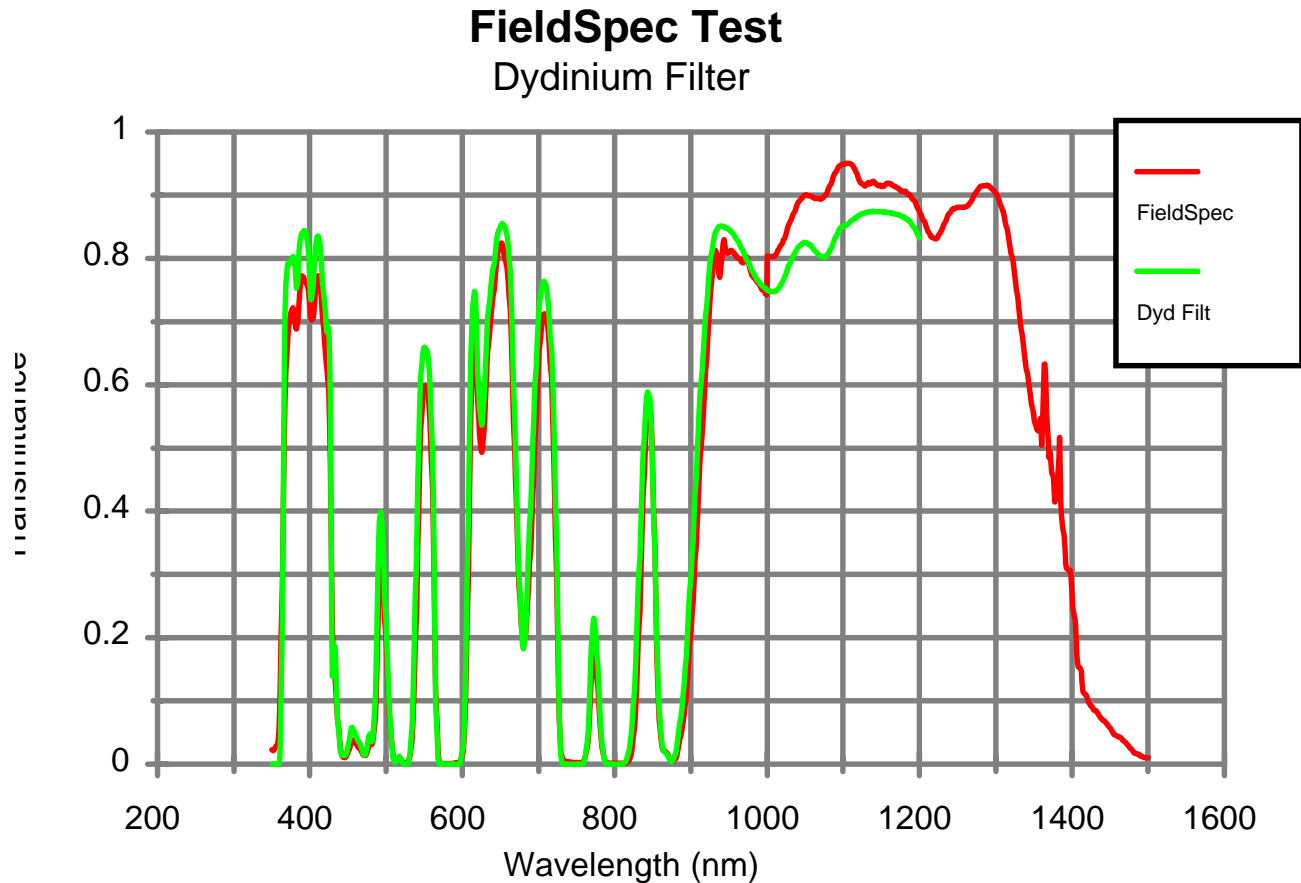
Wavelength Calibration



- ❖ Wavelength calibration is best done by instrument maker
- ❖ Tests for stability should always be done in the field
- ❖ Dydinium filter transmittance using irradiance or panel is a useful method



Testing FieldSpec Wavelength Cal



Measuring Field Spectra using a Reference Panel

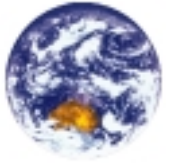


$$\rho_t(\mu_r, E_d) = \frac{\pi L_t(\mu_r)}{E_d}$$

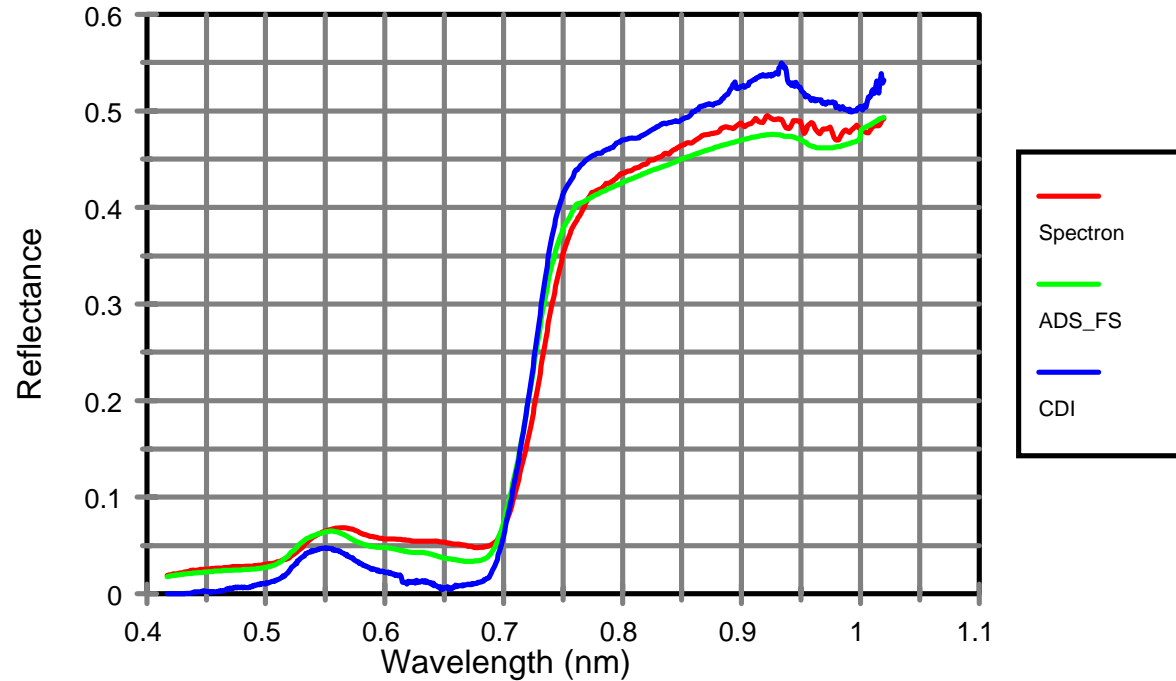
$$\rho_s(\mu_r, E_d) = \frac{\pi L_s(\mu_r)}{E_d}$$

$$\begin{aligned}\rho_t(\mu_r, E_d) &= \rho_s(\mu_r, E_d) \frac{L_t(\mu_r)}{L_s(\mu_r)} \\ &= \rho_s(\mu_r, E_d) \frac{V_t}{V_s}\end{aligned}$$

Grass Spectra at Dickson – The SAME Grass & Panel!



Dickson Oval
Spectroradiometer Comparison



Varying Results for the same target?

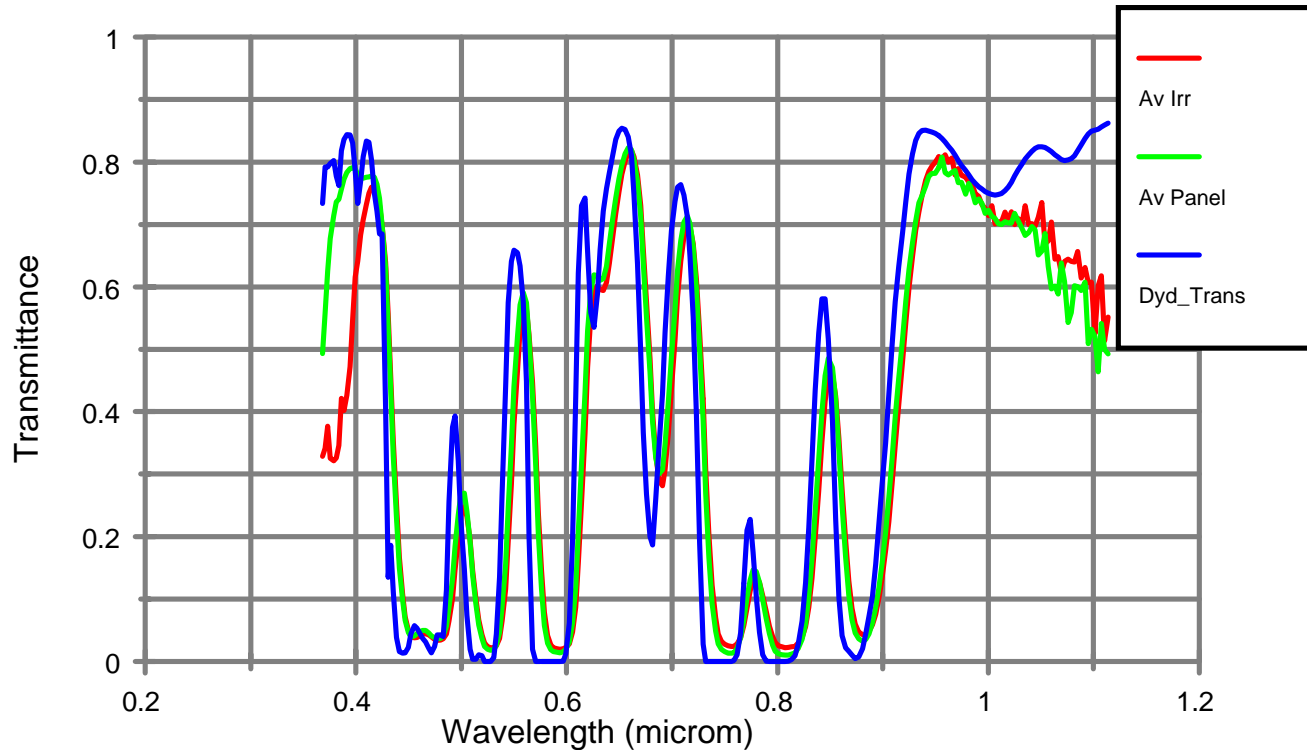


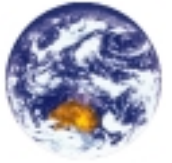
- ❖ The spectron had lost wavelength calibration
- ❖ The CDI detector radiometric calibrations seem to be variable
- ❖ The ASD has unacceptable “jumps” even after division by panel radiance



Testing Spectron Wavelength Cal

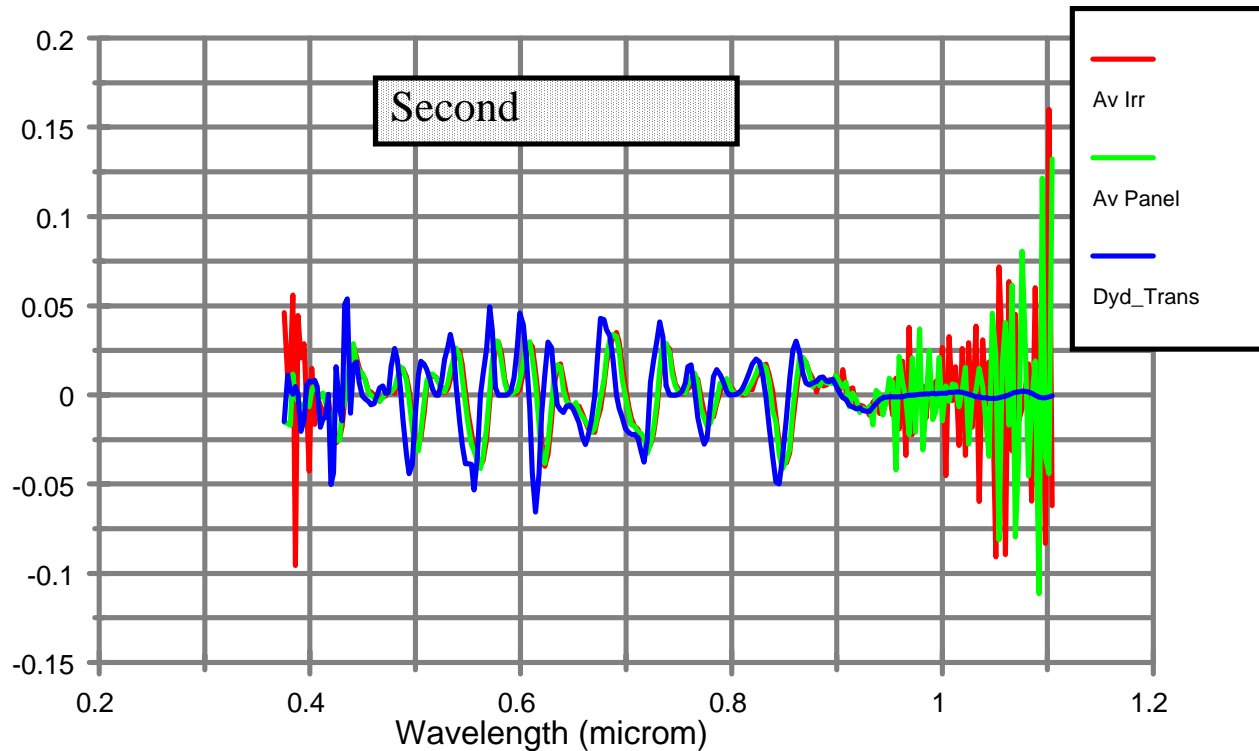
Dydinium Filter
Spectron SE590





Second Derivatives of Spectra

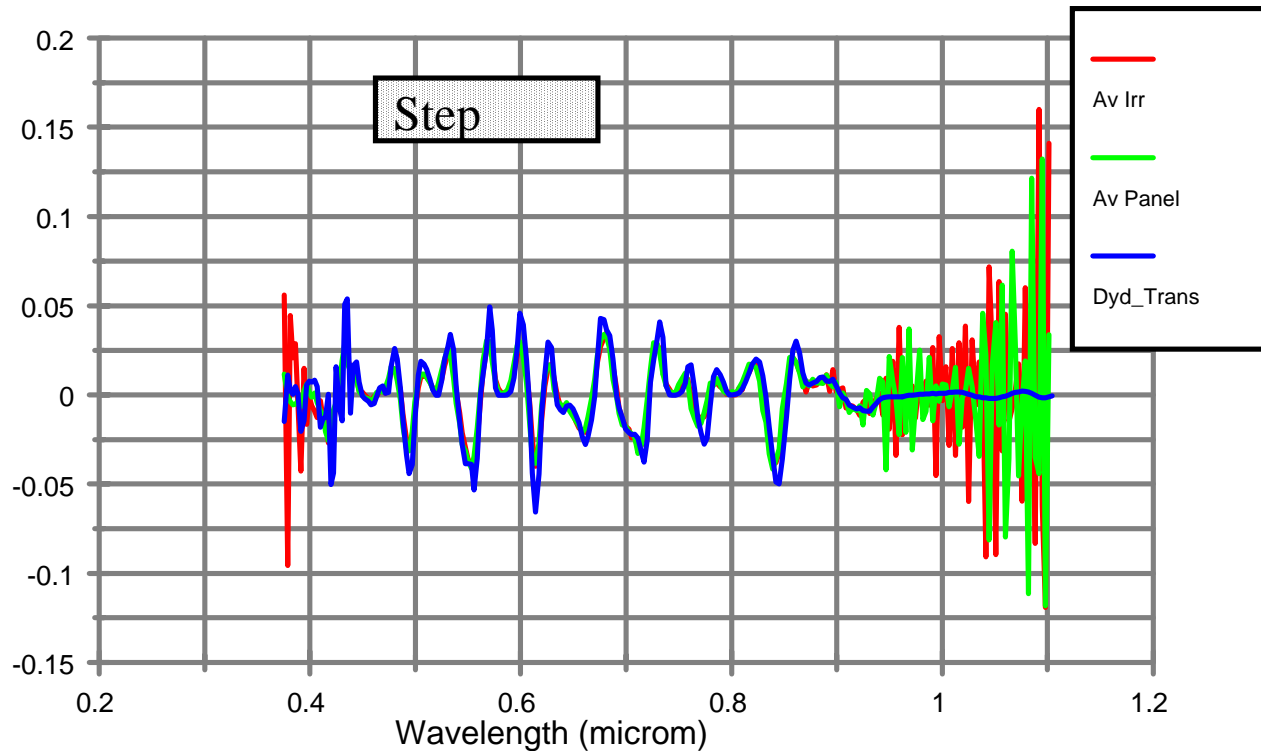
Dydinium Filter
Spectron SE590





Shifted 3 Detectors

Dydinium Filter
Spectron SE590

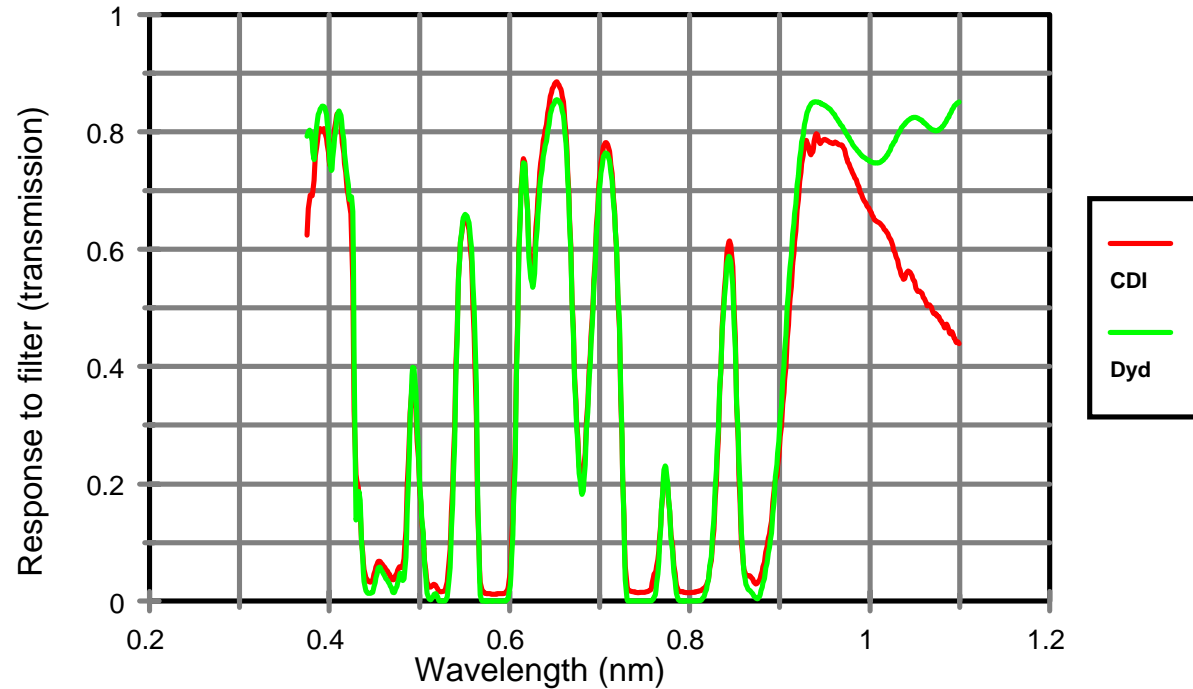


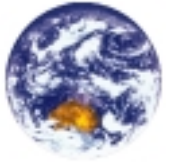


Testing CDI Wavelength Cal

CDI wavelength cal.

BaSO₄ pan. and Dydinium filter- 3-4-98





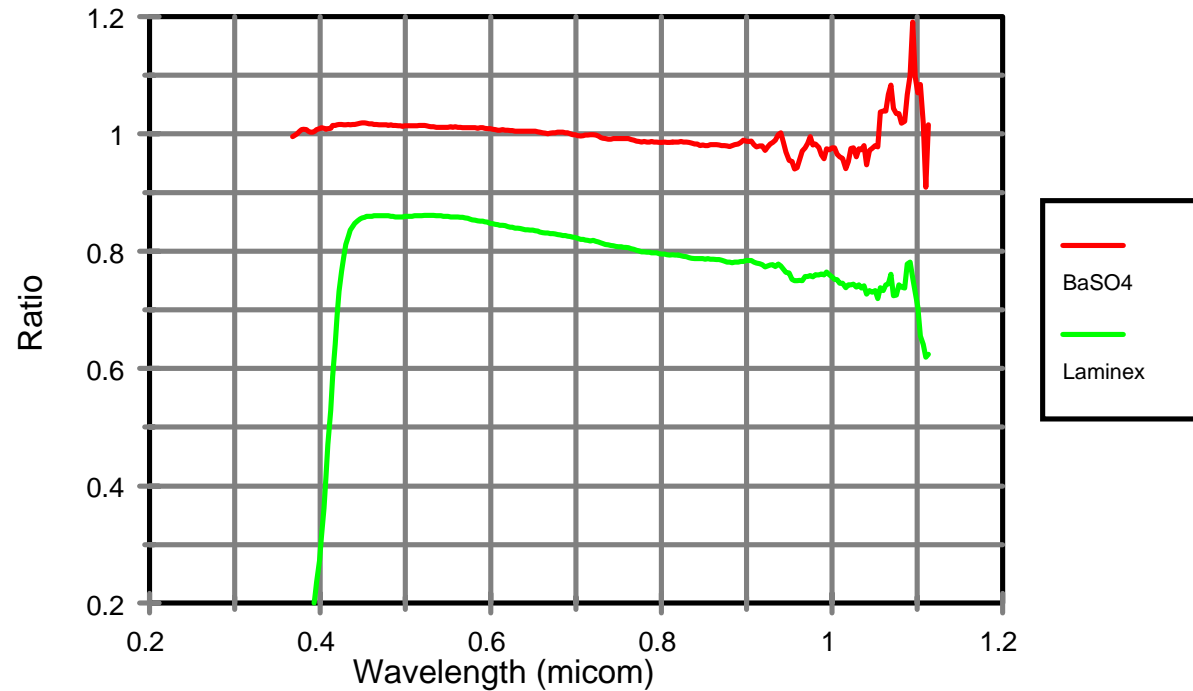
Panel “Calibration”

- ❖ Panels must also be calibrated and tested
- ❖ Laboratory reference panel must be carefully maintained
- ❖ Field standards need testing after each mission
- ❖ Panel K factors – including BRDF must be known



Comparing Panels

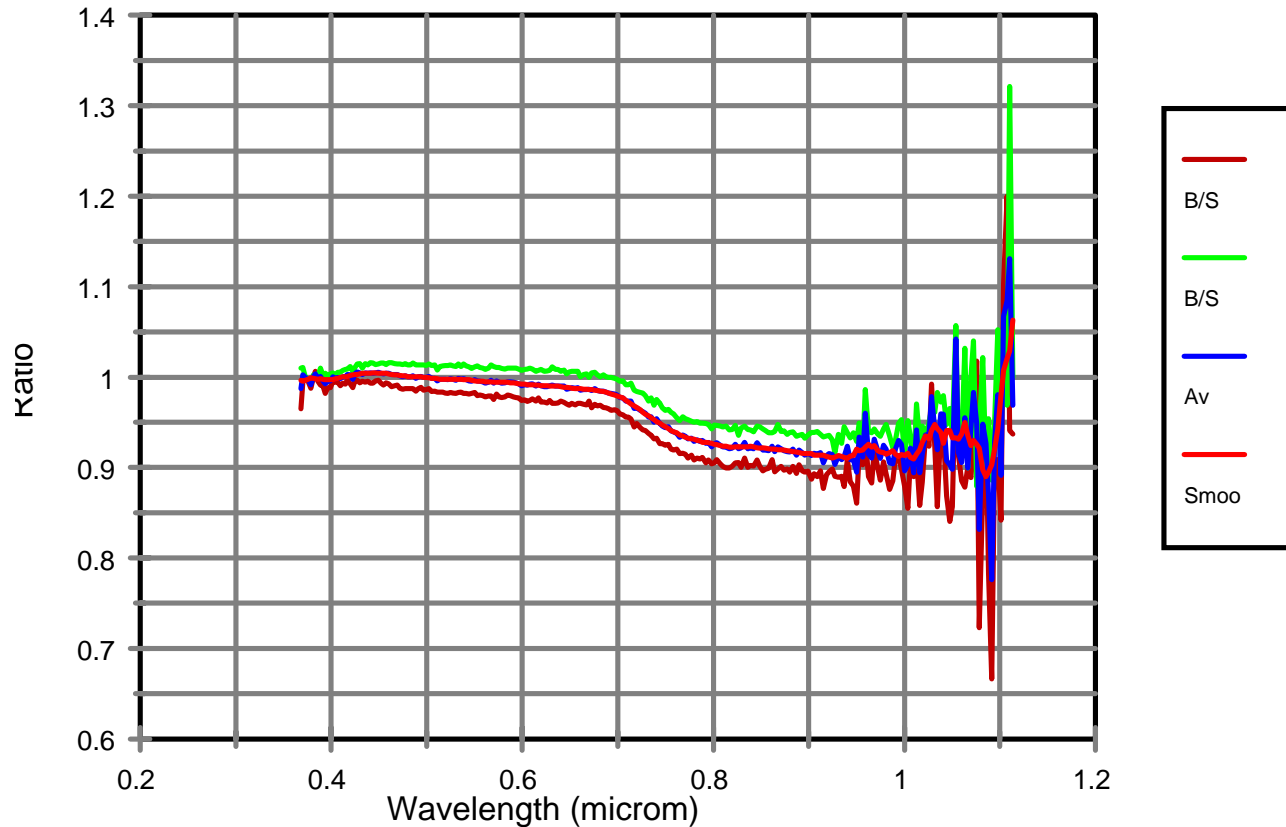
Moreton Panel Check Panel over Spectralon





Panel Variations

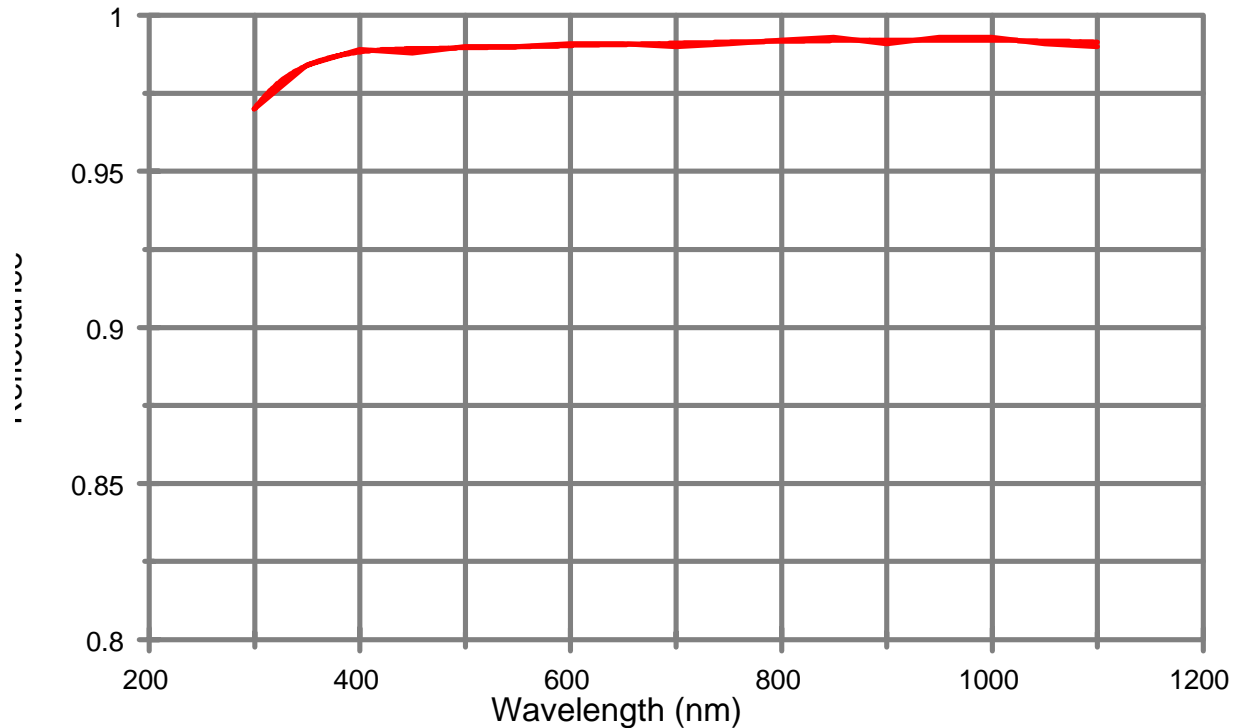
Panel Comparison
Dickson Oval June 1997



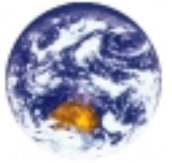
Spectralon Published Reflectance



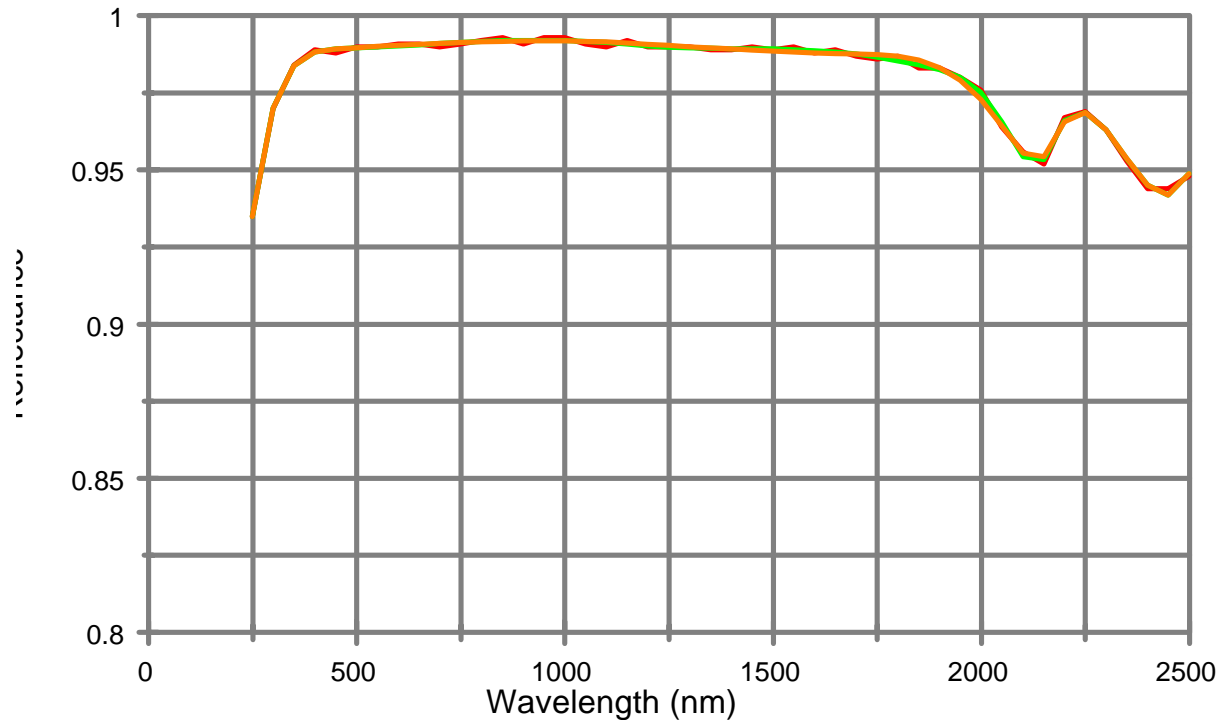
Spectralon Panel
Reflectance



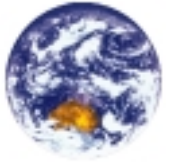
Published Spectralon Reflectance



Spectralon Panel
Reflectance



Using the Data



- ❖ Atmospheric Irradiance Model can be used to set parameters for the atmosphere
- ❖ Field spectra can be used to modify the atmospheric correction of the airborne data
- ❖ Changes in atmospheric conditions can be monitored by Weather Station
- ❖ BRDF effects can be established in the field

Atmospheric Model



$$L_t(\mu_o, \mu_s, h, \lambda) = \frac{1}{\pi} E_T t(\mu_o, h, \lambda) \frac{\rho_t + \rho_{env}}{1 - s\rho^*} + L_p(\mu_o, \mu_s, h, \lambda)$$

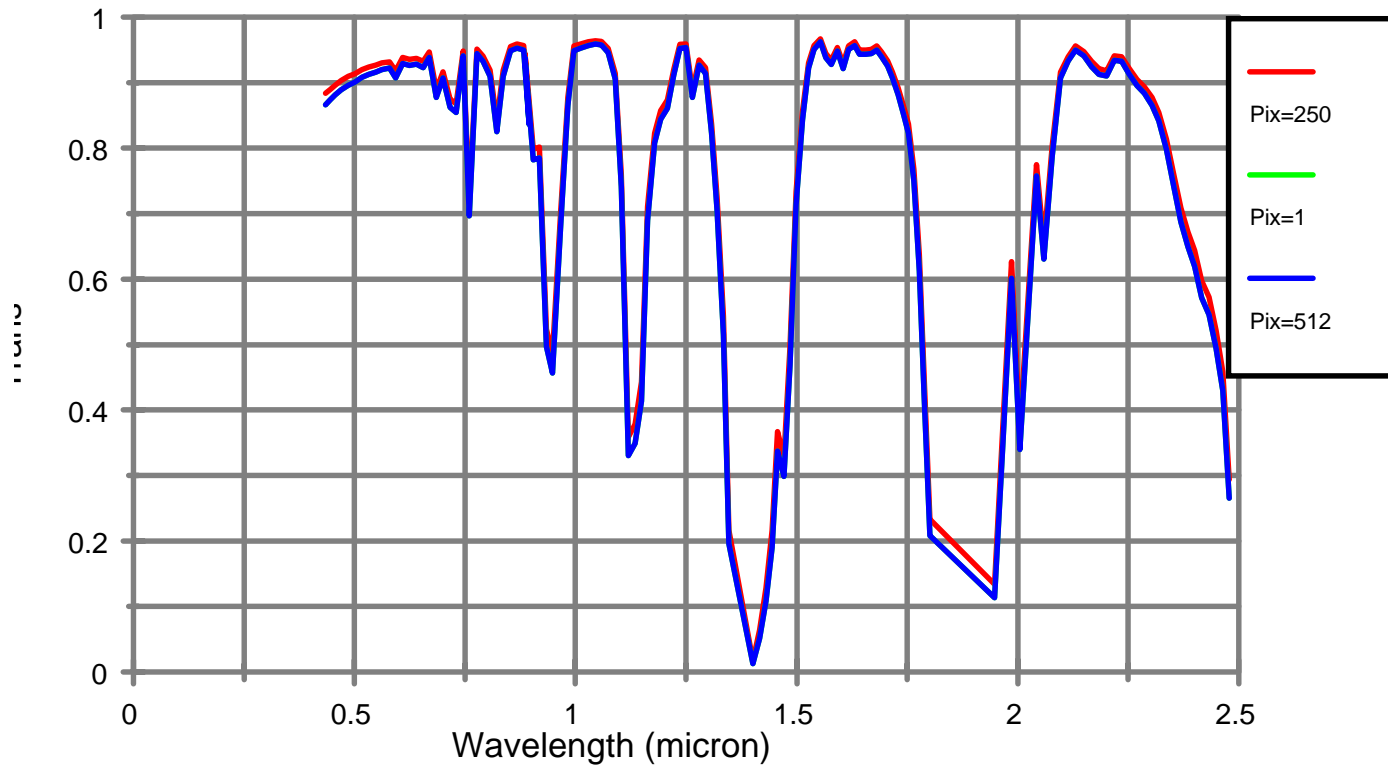
$$\rho_{env} = \rho^* \left[\frac{T(\mu_o, h, \lambda)}{t(\mu_o, h, \lambda)} - 1 \right]$$

$$\left[E_T, t(\mu_o), T(\mu_o), T(\mu_s), s, s(h), L_p(\mu_o) \right], \left[L_g(\mu_o) \right]$$

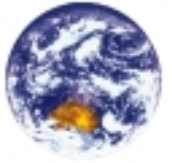
Hymap Simulated Transmittance (aircraft at 10,000 feet)



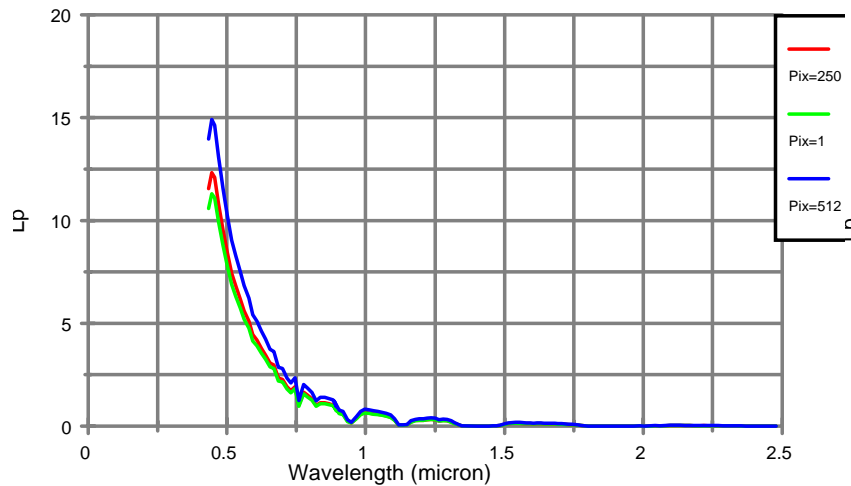
Hymap Run
Transmittance (Heading NW)



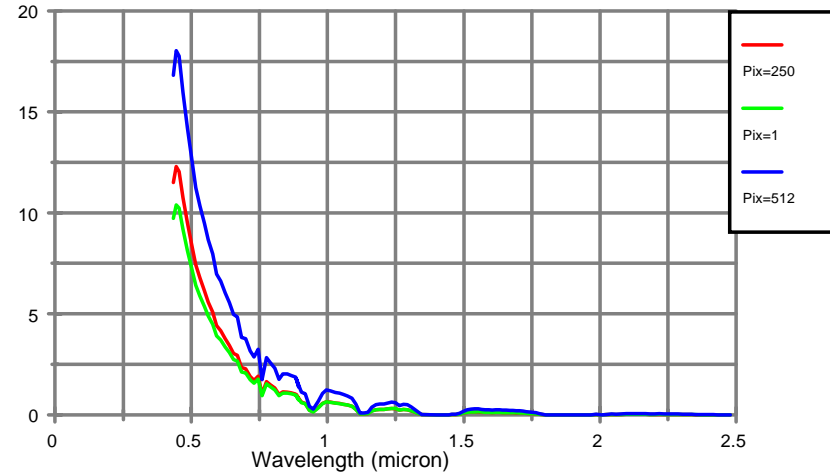
Path Radiance MVA Effects



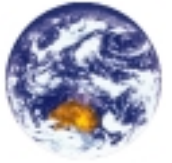
Hymap Run
Path Radiance Heading N



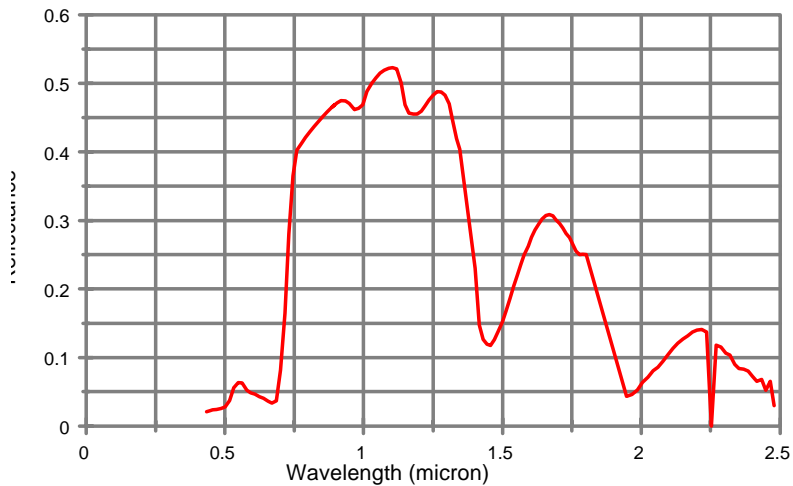
Hymap Run
Path Radiance (Heading NW)



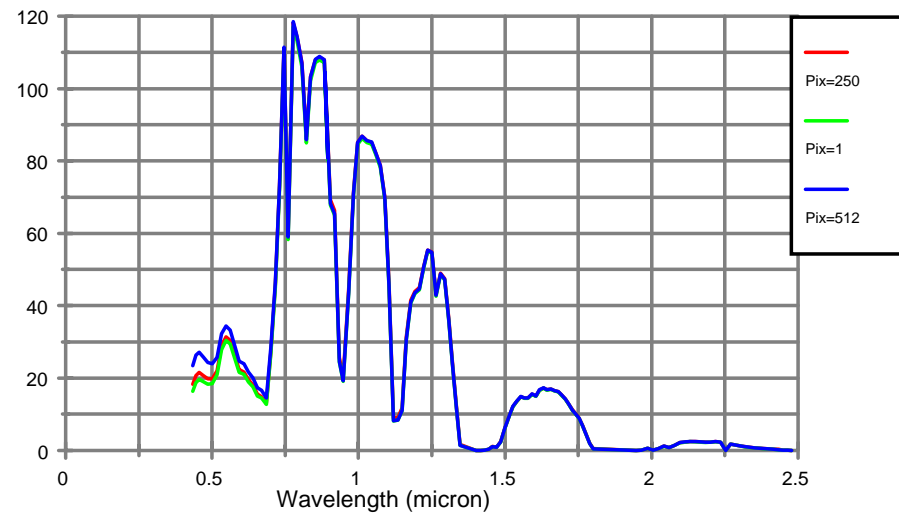
HyMap Simulated Grass Signature



Hymap Run
grass signature



Hymap Run
grass signature



Atmospheric Correction



$$L^*(\mu_o) = \frac{1}{\pi} E_T T(\mu_o) \frac{\rho^*}{1 - s\rho^*} + L_p(\mu_o)$$

$$L_{\min}(\mu_o) = \frac{1}{\pi} E_T (T(\mu_o) - t(\mu_o)) \frac{\rho^*}{1 - s\rho^*} + L_p(\mu_o)$$

$$\rho_t = \frac{\pi(1 - s\rho^*) \left(L_t - \frac{L_p(\mu_o)}{(1 - s(h)\rho^*)} [-L_g(\mu_o)] \right)}{E_T t(\mu_o)} - \rho_{env}$$